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Key Questions

- 1. What instrumentation projects to put forward for NSF funding or small upgrades funded by CARA?
- 2. Should UCOAC endorse Keck AOAG recommendation to move forward with NGAO?
- 3. Should UC and Caltech re-examine their partnership agreement now to enable improved institutional fund-raising for WMKO?
- 4. How can UC help to ensure that UC-WMKO instrumentation is successful? (& assure their partnership share?)

WMKO Projects that Have Been Approved for NSF Proposal

- KCWI –red
 - PI = C. Martin (Caltech)
 - Estimated cost = \$2.8M
 - Funding Target = NSF-MRI
 - Unsuccessful 2012 (passed PDR, split from KCWI-blue to finish under TSIP)
 - Plan resubmit Jan 2013
- OSIRIS Detector Upgrade
 - PI = J. Larkin (UCLA)
 - Estimated cost = \$1M
 - Funding Target = NSF-ATI (WMKO?)
 - Plan to submit to NSF-ATI this round
- K1DM3
 - PI = X. Prochaska (UCSC)
 - Estimated cost = \$1.7M
 - Funding Target = NSF-MRI
 - Unsuccessful 2012
 - Resubmit?

NGAO – Well Advanced But Funding Limited

- NGAO
 - Passed Preliminary Design Review in June 2010
 - Estimated remaining cost
 - AO = \$37M
 - DAVINCI = \$13M
 - Board placed hold on until funding path clear
 - New Opportunities on Horizon
 - Mid-scale recommendation Astro2010 & NSF AST Portfolio Review (Aug 2012)
 - Keck 20th Anniversary Celebration
 - Keck AO Advisory Group
 - Keck SSC formed this sub-committee this summer
 - Is NGAO still long-term AO strategy?
 - Recommendation Yes, move forward.
 - UCLA Development office strongly engaged
 - UCOAC endorsement?

Issues

- Original WMKO Operational Model run out of gas before
 2018
 - Key decision in 2005 that effected this
 - Instruments Originally funded Entirely by CARA
 - LRIS-red (1993; Caltech)
 - HIRES (1994; UCSC)
 - NIRC (1994; Caltech)
 - LRIS-blue (2000; Caltech)
 - LWS (1996; UCSD, UCB)
 - HIRES image rotator (1997; UCSD)
 - KCAM Modification (1998; UCLA)
 - ESI (1999; UCSC)
 - NIRSPEC (1999; UCLA)
 - K2 AO [NGS] (1999)
 - NIRSPEC AO Upgrade, NIRSPAO (2000)
 - NIRC2 (2001)
 - K1 NGS-AO (2001)
 - DEIMOS (2002)
 - K2 LGS-AO (2004)
 - Interferometer (2004)
 - HIRES Detector Upgrade (2004)

Issues

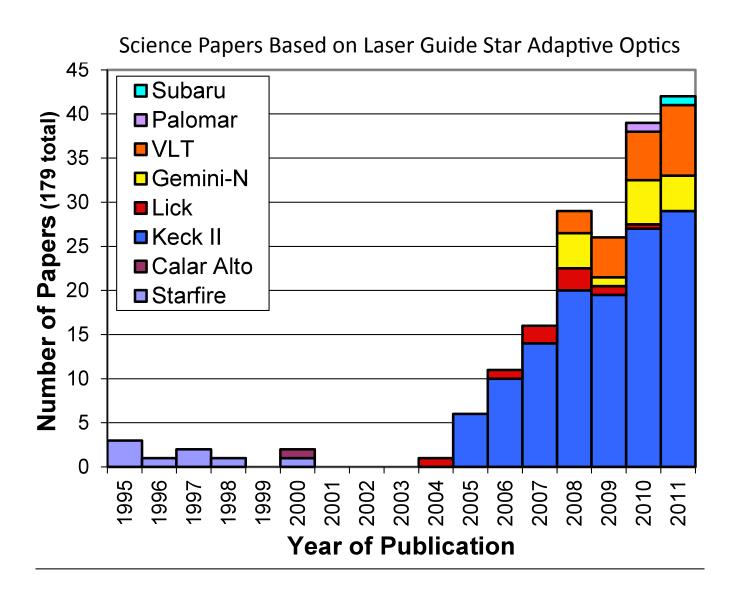
- Original WMKO Operational Model run out of gas before 2018
 - Instruments funded Entirely with some or all TSIP
 - OSIRIS (2005)
 - MOSFIRE (2012)
 - KCWI-blue (TBD)
 - Instruments funded Entirely outside TSIP
 - K2 Wave Front Controller Upgrade (2007) Keck Foundation
 - K1 AO Upgrade (TRICK) NSF-ATI
 - K2 AO Upgrade (Center Launch) NSF-MRI
 - K2 AO Upgrade (Laser) Keck Foundation, Moore Foundation & Friends of Keck
 - Mirror Segment Microfractures proposed for current night "exchange"

UCOAC Input

- NGAO go-ahead
 - For mid-scale?
 - For Gala vision and fund-raising?
 - Engage UC development?
 - NGAO Observatory Tour at Gala? Other Ideas?
 - If so, does UC gain nights?
- K1DM3 resubmission
 - Expanded TDA operational model?
 - Straw-man expansion
 - Multi-institution ToO program
 - Cadence observing (up to 1hr per night)
 - Optimize observing to conditions (TAC-approved bad-seeing programs)
 - Expanding these activities may mean reducing other operations

Slides from AOAG presentation at Keck Science Meeting

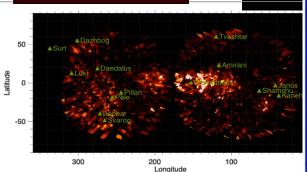
Keck Leads in the field of Adaptive Optics



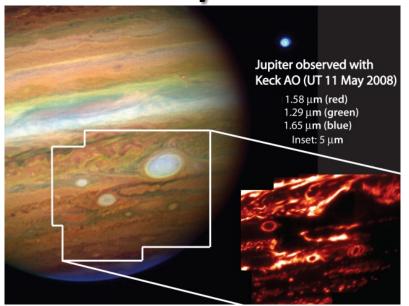


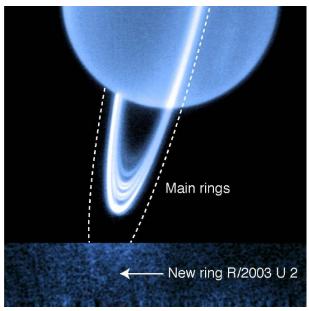


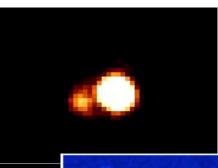


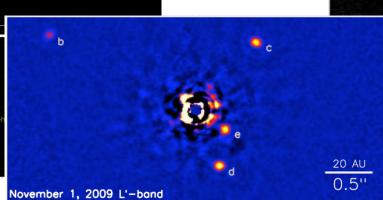


Solar System Science





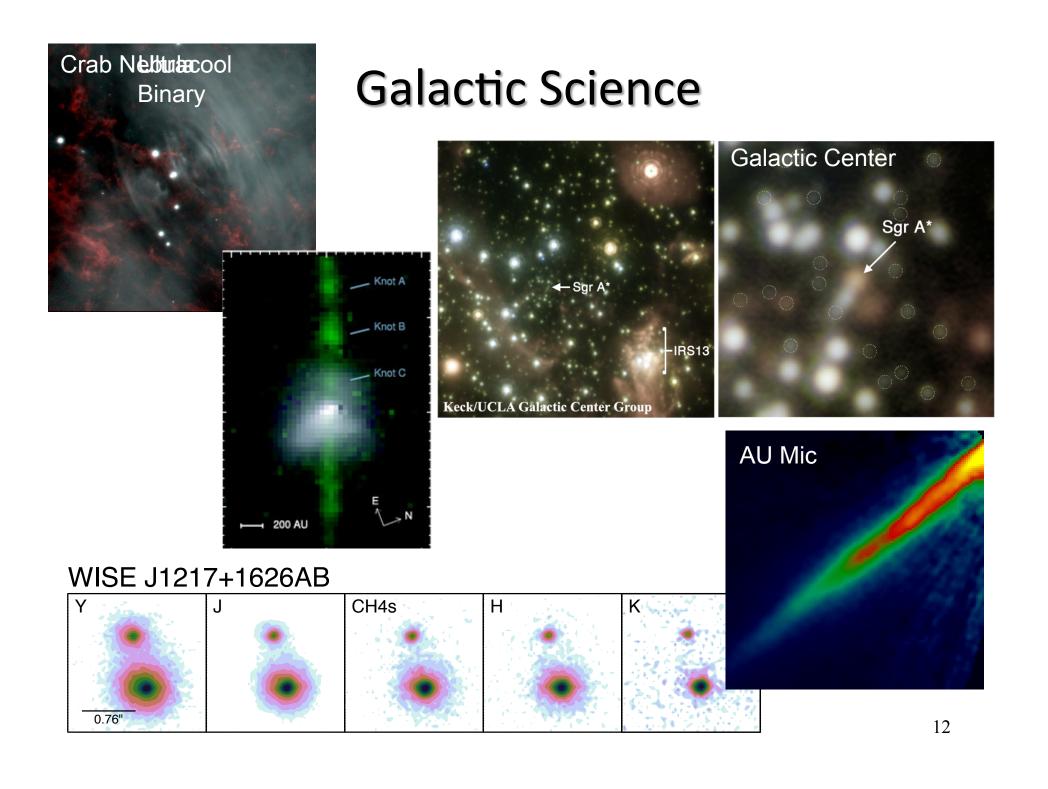




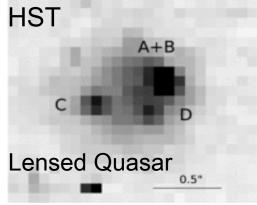


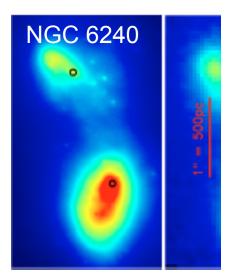
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Nix and Hydra

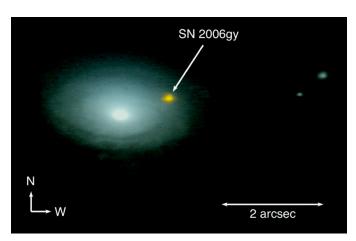


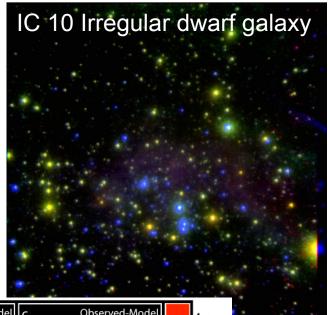
Keck LGS B A C D 0.5"

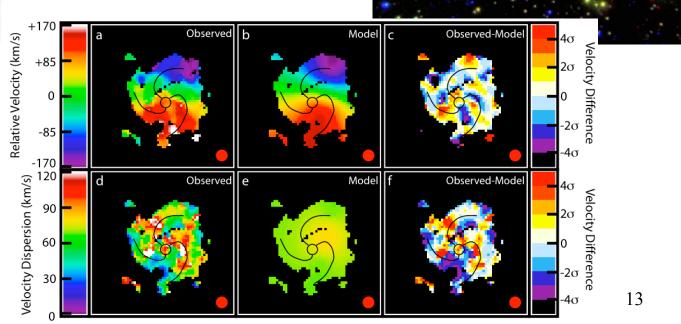




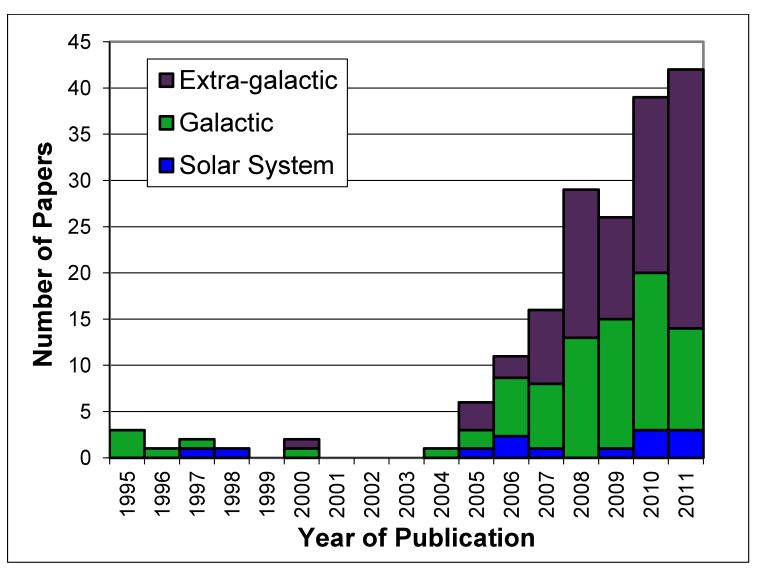
Extragalactic Science





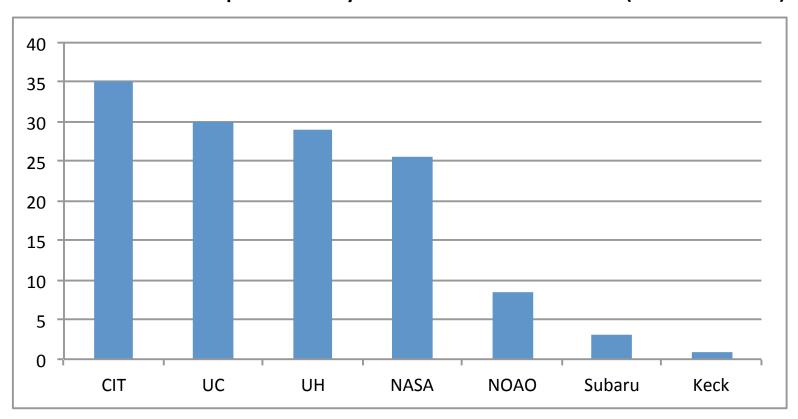


Rapid Growth in Extra-Galactic AO Science Since Advent of LGS

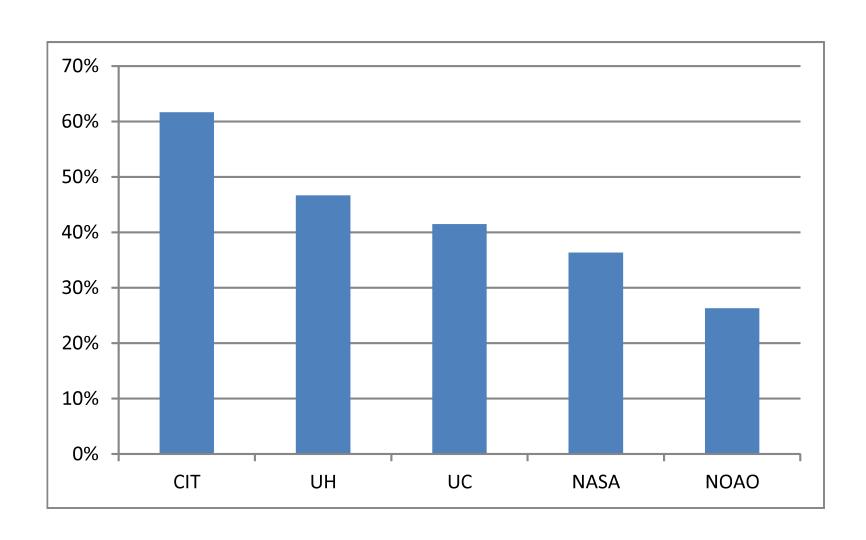


Broad Usage Across Community

Number of Unique PI's by Partner Institution (2008-2012)



AO PIs as a % of all Keck PIs in 2011

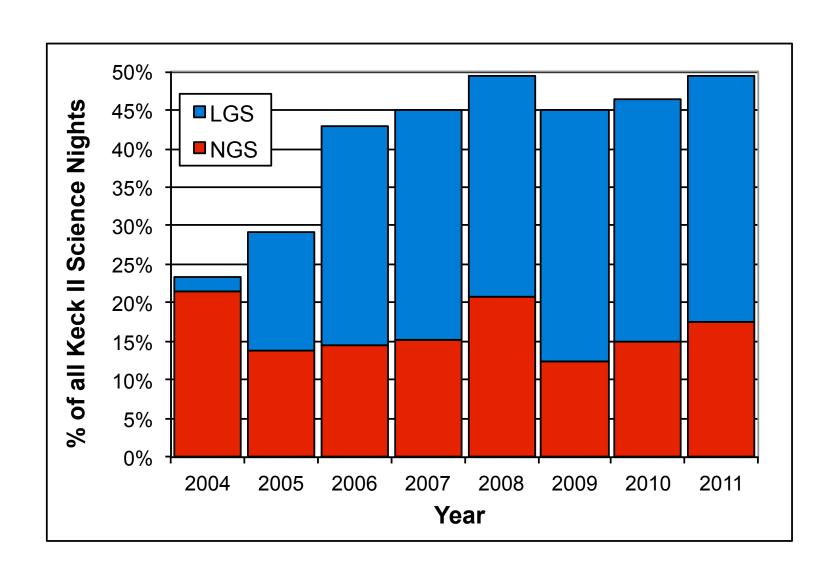


High Demand by Graduate Students

- Of the 43 PhD awarded 2009-2010 based on Keck data
 - At least 42% of these contain Keck-AO data
 - 49% of these students appeared as co-authors on papers with Keck-AO
 - 90% of those that won Prize Postdoc Fellowships used
 AO in their thesis (9/10)



Number of AO Award Nights is High



Keck AO has Already had Tremendous Science Impact

but

Full Potential of AO Has Not Yet Been Achieved

Next-Generation Adaptive Optics (NGAO) Primer



Peter Wizinowich Principal Investigator



Claire Max Project Scientist

Capabilities Based on Broad Science Reach

Key Science Goals

Understanding the Formation and Evolution of Today's Galaxies since z=3

Measuring Dark Matter in our Galaxy and Beyond

Testing the Theory of General Relativity in the Galactic Center

Understanding the Formation of Planetary Systems around Nearby Stars

Exploring the Origins of Our Solar System

Key New Science Capabilities

Near Diffraction-Limited in Near-IR (K-Strehl ~80%)

AO correction at Red Wavelengths (0.7-1.0 μm)

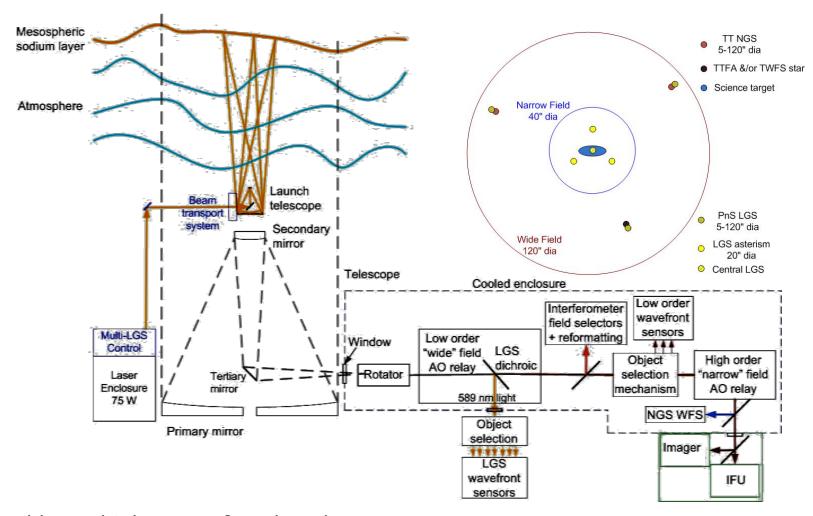
Increased Sky Coverage

Improved Angular Resolution, Sensitivity and Contrast

Improved Photometric and Astrometric Accuracy

Imaging and Integral Field Spectroscopy

NGAO System Architecture



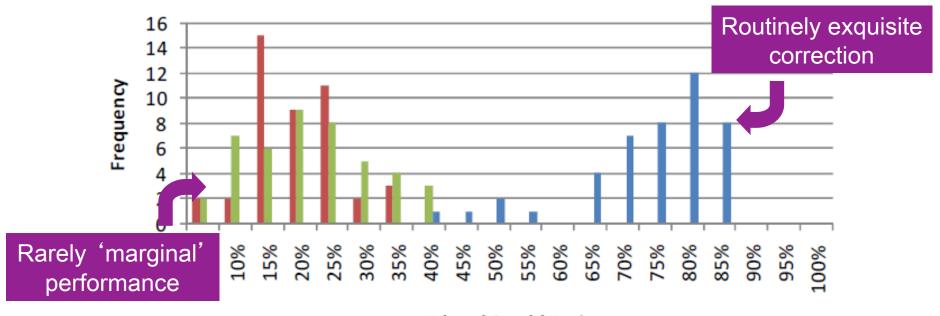
High Strehl = Multiple Lasers & High Order DM

High Sky-Coverage = AO-corrected NIR Tip-Tilt sensor & Multiple-Tip/Tilt stars

High Sensitivity = above + cooled enclosure

NGAO will change the AO observing experience

AO performance comparison



K-band Strehl Ratio

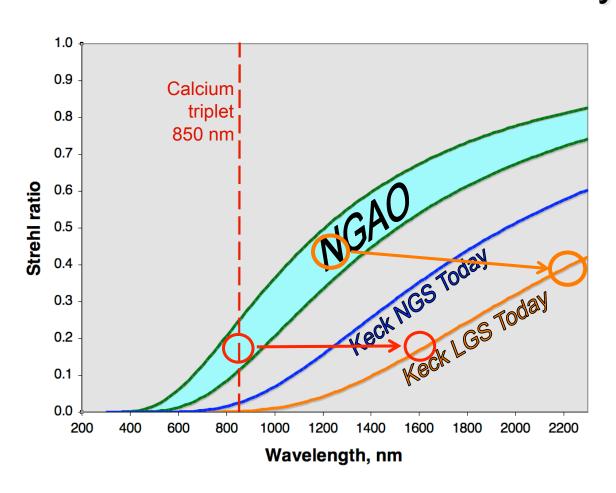
■ Liu's K2 LGS AO Data
■ K2 LGS AO Model
■ NGAO Model

Monte Carlo performance estimate simulating 44 nights observing (Galaxy Assembly science case), drawing random values for r₀, wind speed, sodium abundance, and zenith angle (KAON 716, Figure 12)

Includes comparison with M. Liu's measured K2 LGS data ($\langle SR \rangle = 17\%$), the KAON 721 predict for K2 LGS ($\langle SR \rangle = 20\%$), and NGAO predict ($\langle SR \rangle = 70\%$)

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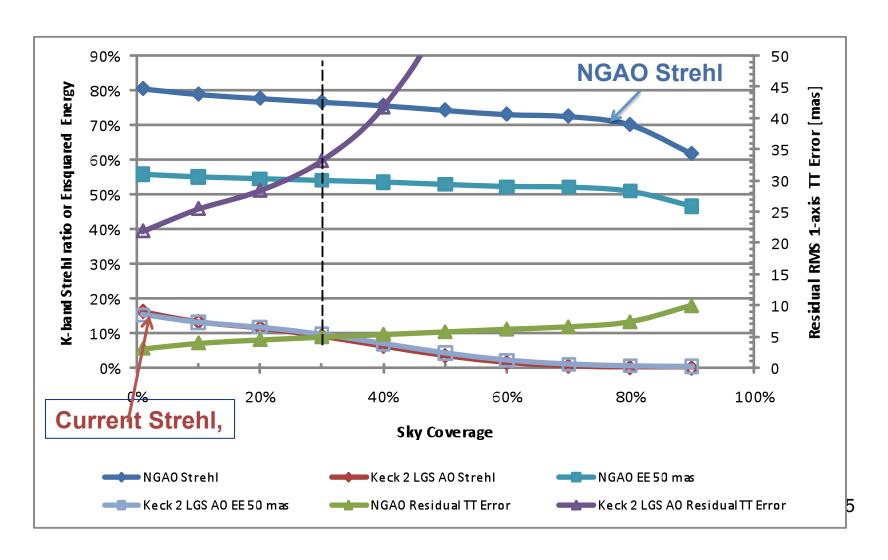
How is NGAO different from Keck's AO today?



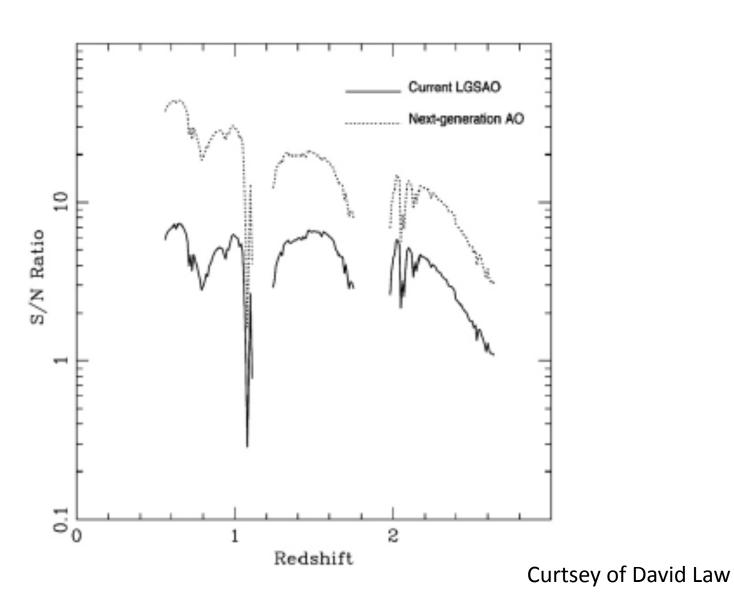
- NGAO: same
 Strehl at Ca
 Triplet as LGS
 today in H band
- NGAO: better
 Strehl at J band
 than LGS today in
 K band
- NGAO: much higher sky coverage

Sky coverage of NGAO is dramatically better than Keck 2 LGS AO

Galaxy Assembly science case (b=60)
Median seeing, sodium return a bit worse than average



Dramatic Improvement in Sensitivity

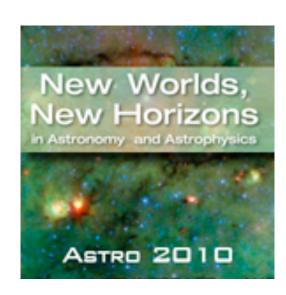


AOAG Early Assessment

- NGAO remains a flagship for WMKO
 - Broad and important science (strehl & sky coverage)
 - Increasing interactions with TMT
 - Key to our future success as a world-class observatory
- New Fund-Raising Opportunities
 - Mid-Scale program at NSF
 - Keck 20th Gala & Science Celebration

Mid-Scale Innovations Program

- Recommended by Decadal Survey (Astro2010) August 2010
 - NGAO was given as an examplar (1 of 7)
- Reinforced by NSF Astronomy Portfolio Review August 2012
- Acknowledged by NSF Astronomy Response August 2012
- NSF evaluating Mid-scale program across Math and Physical Sciences (MPS) Now
- NGAO well-positioned to propose
 - PDR passed, including a good cost estimate
 - Since PDR, significant risk reduction underway.



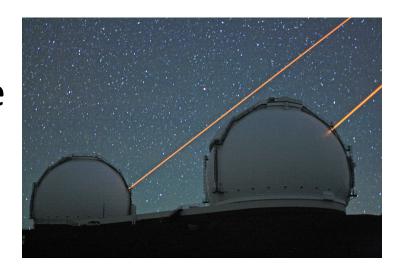


Keck 20th Gala a Great Opportunity

- Clearly communicate to broad audience that NGAO is Keck's Flagship
 - Institutional leadership
 - Foundations (increasing interest in AO)
 - Donors
 - Federal Funding Agencies
- Create a community of financial supporters
 - Existing Friends of Keck have significant resources
 - Reach new potential donors
 - Friends of Keck network
 - Friends of Partners
 - UC inviting 150 potential donors
 - UCLA is already actively engaged

Next Steps for AOAG

 Prepare for NSF proposal opportunity for mid-scale program (e.g., update science case, re-evaluate competitive landscape)



Create a private fund-raising strategy

 Assess benefits & priority of additional staged implementation steps and AO improvements